

Status and Perspectives for Wave Height estimation from altimeter measurements

***P.Thibaut, F.Piras, T.Moreau, L.Amarouche, P.Rieu, (CLS)
F.Boy, N.Picot (CNES)
F.Borde, C.Mavrocordatos (ESA)***

Rationale

- ❑ For many years, most of research is conducted on SLA and debates are mostly focused on it, in particular in this splinter when speaking about retracking
- ❑ But, wave height is also a **key parameter** derived from altimetry

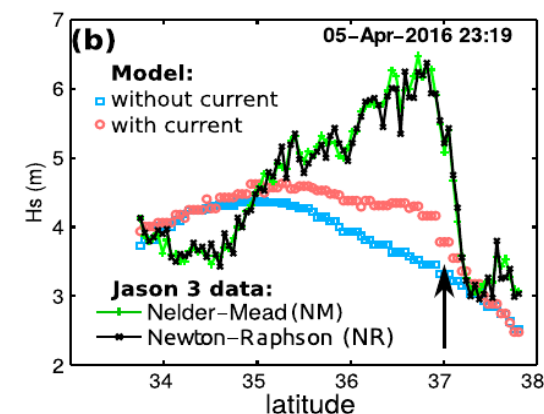
- ❖ for long term climate monitoring
- ❖ for operational marine services
- ❖ for coastal studies (coastal erosion ...)
- ❖ for extreme events
- ❖ because of impacts on SSH via SSB
- ❖ because of its correlation with currents
- ❖ ...

- ❑ Objective: review of wave heights as provided by the different altimeter systems

- ❖ Accuracy
- ❖ Precision
- ❖ Bias/Noise correlated to the wave period

➔ Objective to get higher performances and continuity between missions

Missions are already providing very nice SWH estimates but there are still **areas for improvement** in the processing

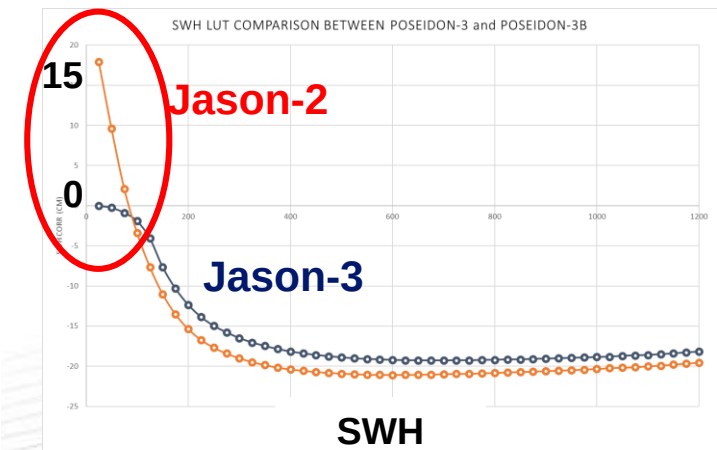
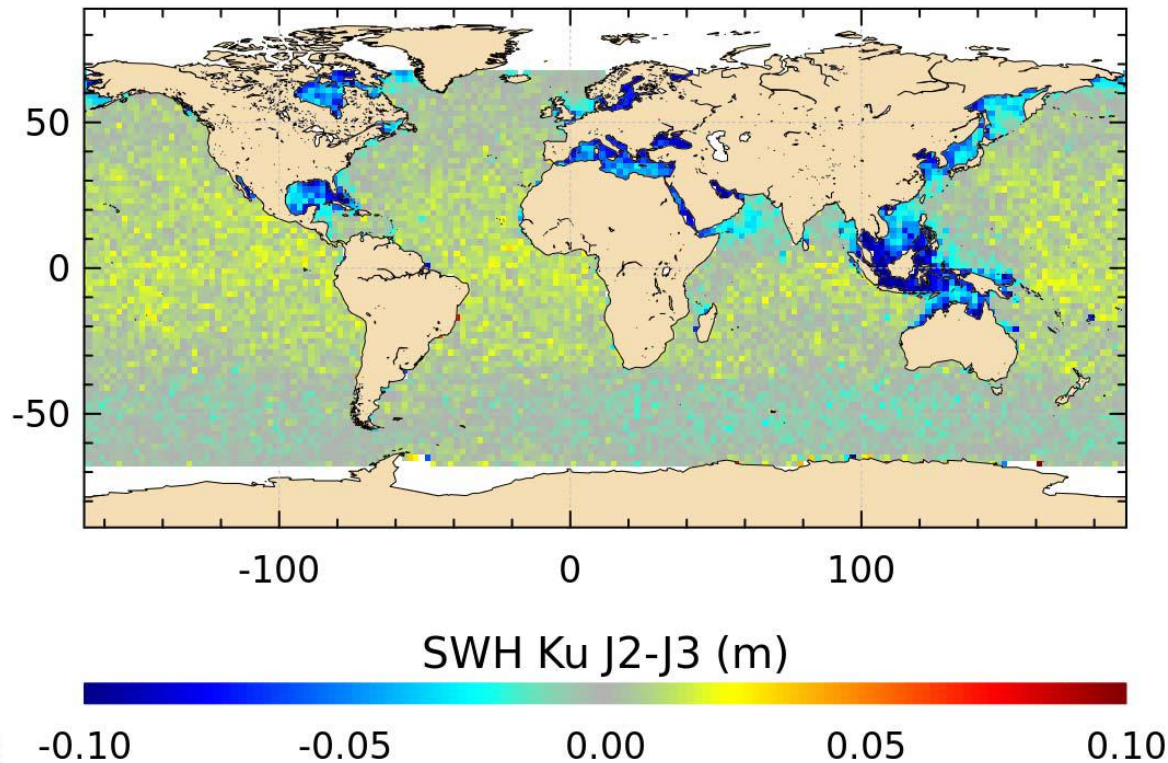


Courtesy F.Ardhuin, JGR 2017

Biases

Jason-2 / Jason-3

- ❑ Jason-2 and Jason-3 fully consistent ($\pm 2\text{cm}$) except for very small SWH (below 1m)
- ❑ Differences are explained by the evolution of LUT corrections (Jason-2 not yet reprocessed)



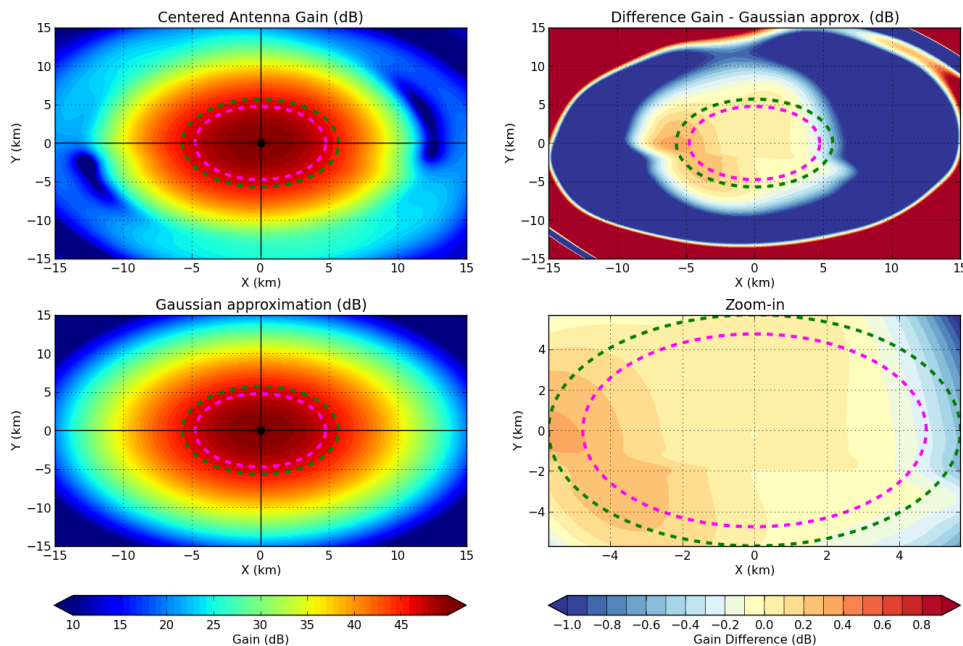
SWH corrections
for POS3 and POS3-B

Biases

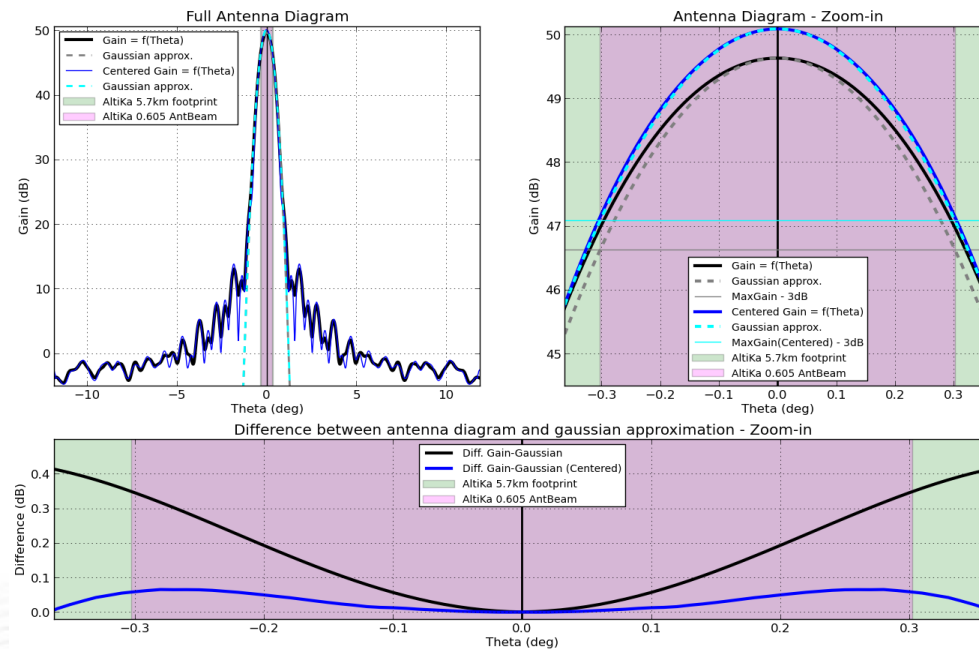
Saral / AltiKa

The gaussian approximation for the Antenna Gain Pattern is not fully valid for small apertures (~ 0.6 deg) (cf S.Le Gac at OSTST, 2016)

AltiKa antenna diagram - Gain=f(X,Y)



AltiKa antenna diagram - Mean Gain=f(Theta)



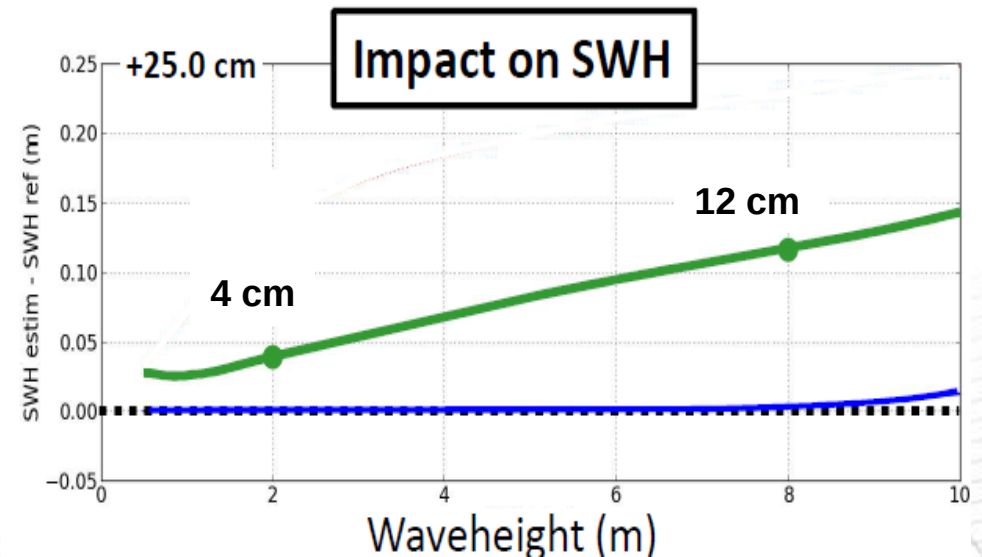
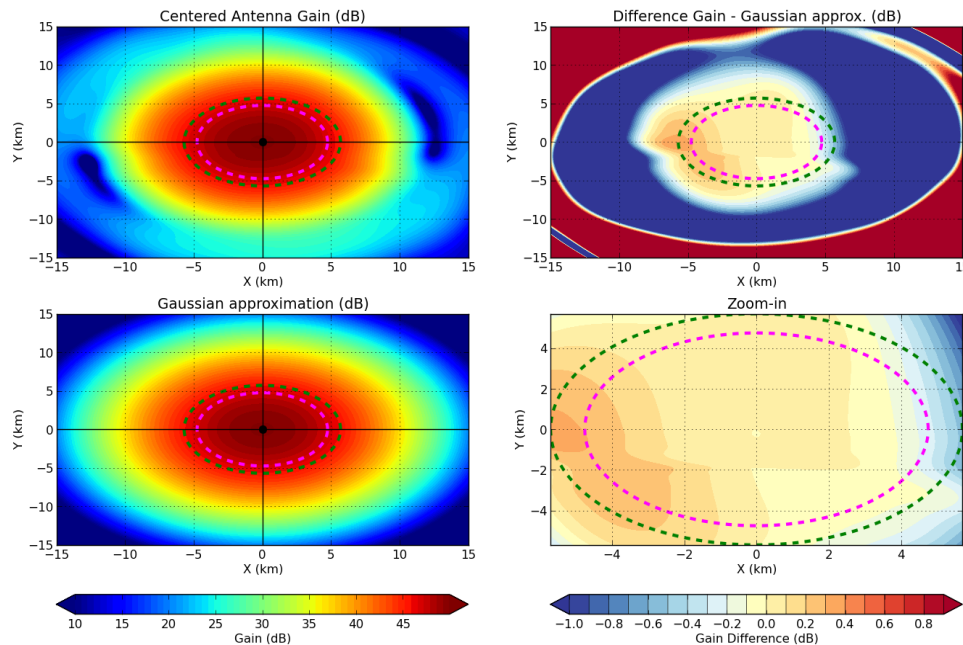
Impacts on SWH can be accounted for by LUT provided as a function of SWH.

Biases

Saral / AltiKa

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AltiKa antenna diagram - Gain=f(X,Y)



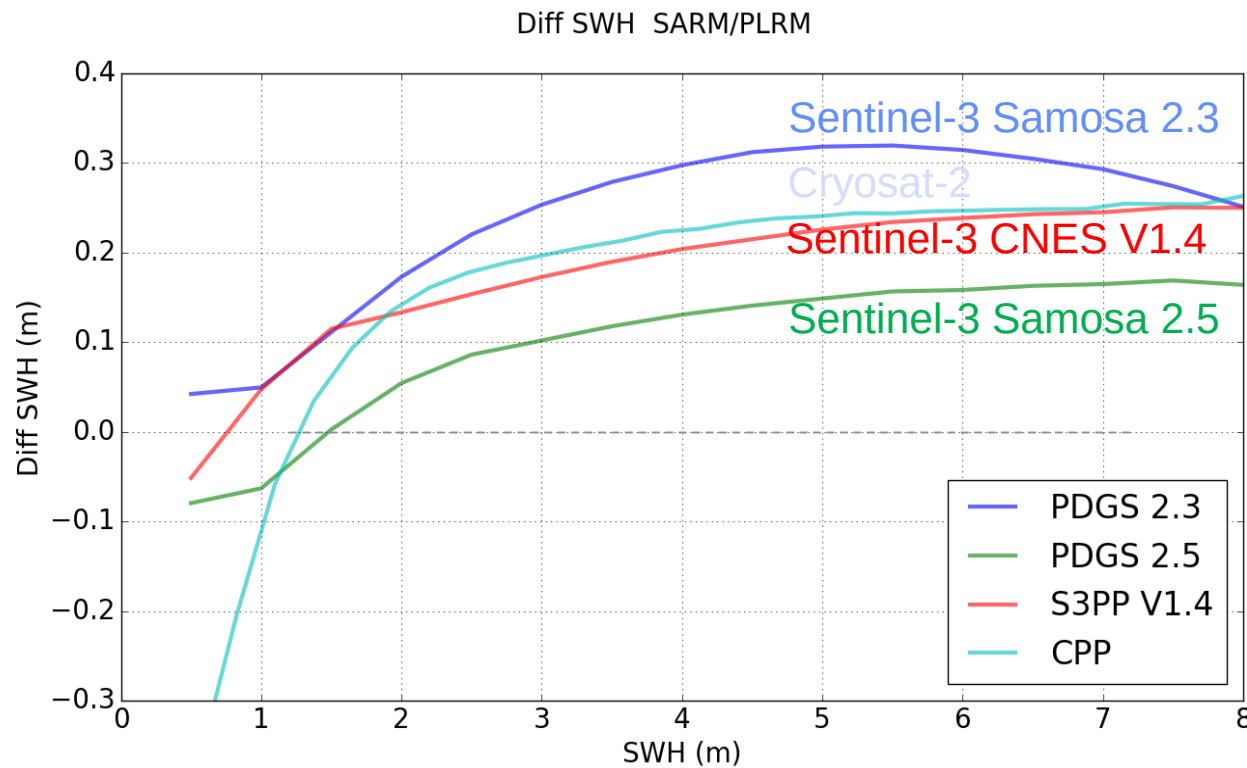
It depends also on pointing angle (gaussian approx. even worse during mispointing events) and we all know that Saral experienced in the past some pointing issues. Ongoing reprocessing activity (GDR-E) at CNES with new LUT including real antenna pattern. Products will be available soon.

Biases

Sentinel-3A

Different SAR processors are showing a similar behavior when looking at SARM/PLRM differences

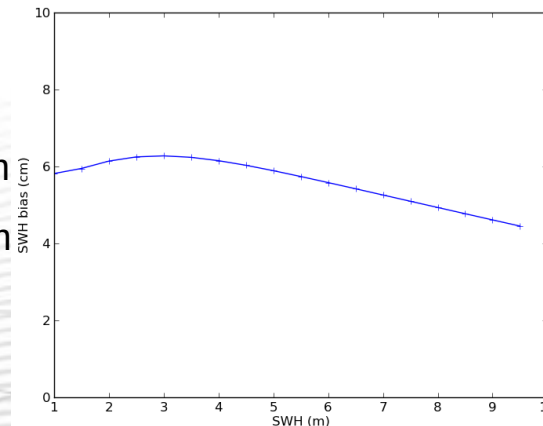
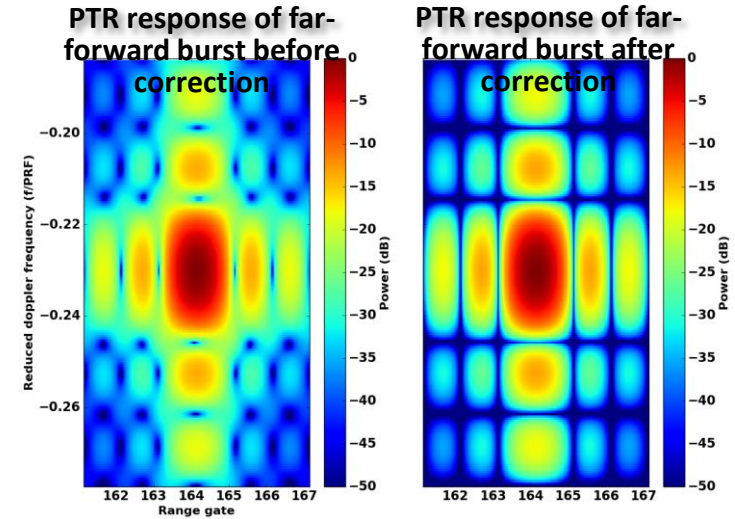
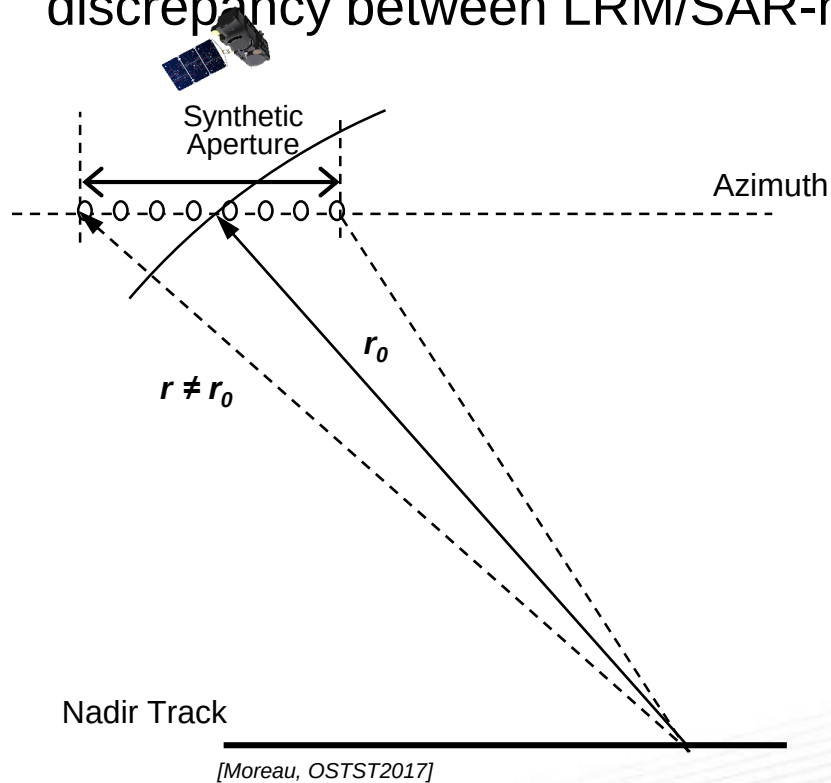
- ❑ ~ bias of about 15cm for SWH > 2m with
- ❑ dependency with SWH in particular for small waves



Biases

Sentinel-3A

Accounting for « range walk » in SAR processing allows reducing SWH discrepancy between LRM/SAR-mode



RW correction removes 5 cm to current SAR SWH with no big dependency with SWH

(see T.Moreau, OSTST 2017)

But not introduced yet in the operational processor (only in CNES processing)

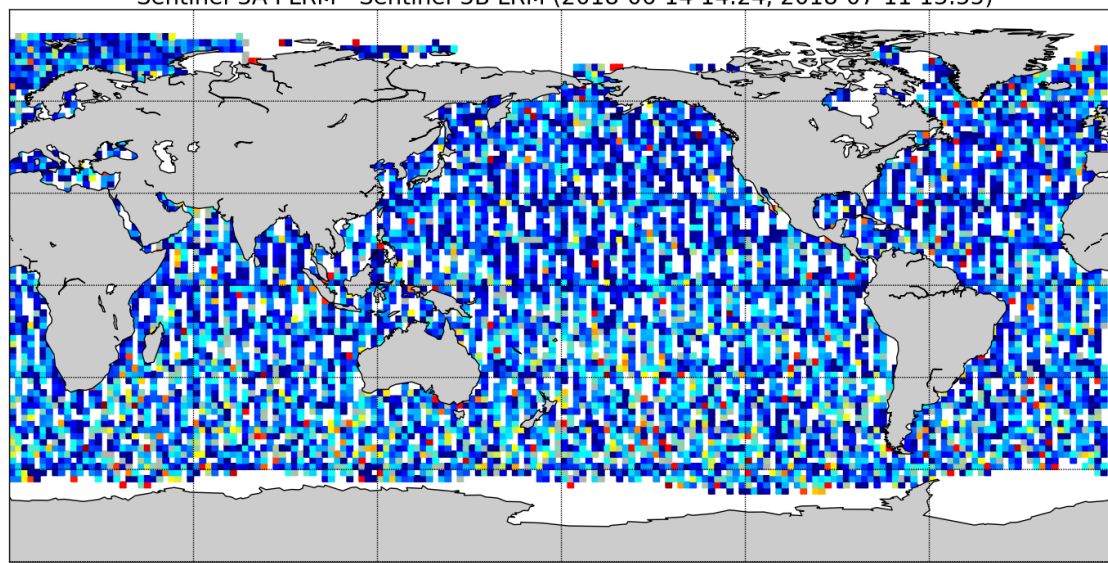
Biases

First look at Sentinel-3A / Sentinel-3B tandem phase

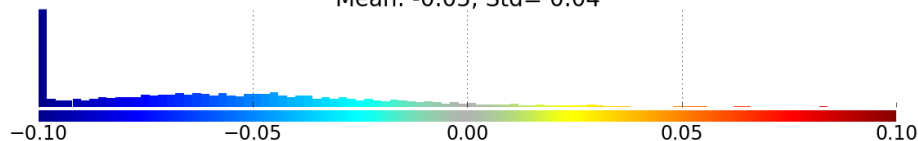
A bias can be seen between PLRM (which is the reference when looking at SAR data) and LRM which is not explained by LUT differences → ongoing activity to better understand these differences

Map of PLRM / LRM SWH differences

Sentinel-3A PLRM - Sentinel-3B LRM (2018-06-14 14:24, 2018-07-11 13:55)

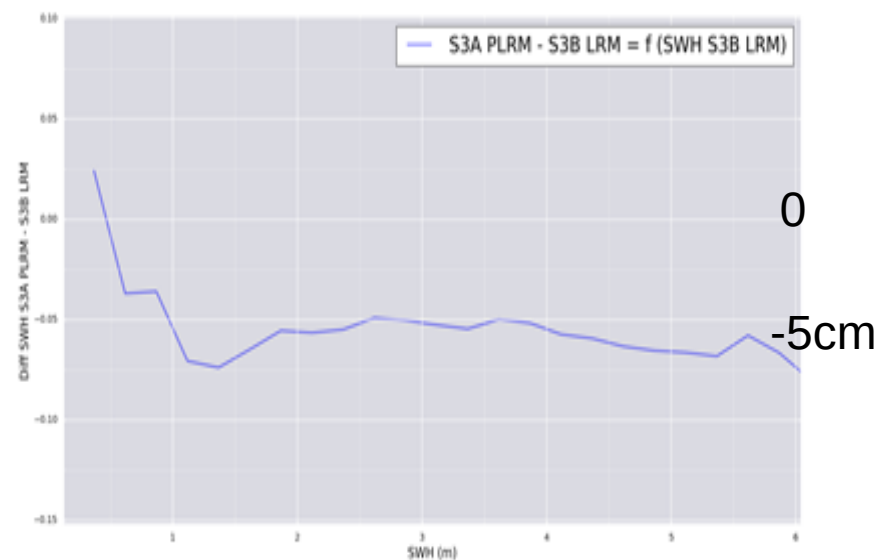


Mean: -0.05, Std= 0.04



(source S3PP CNES processing)

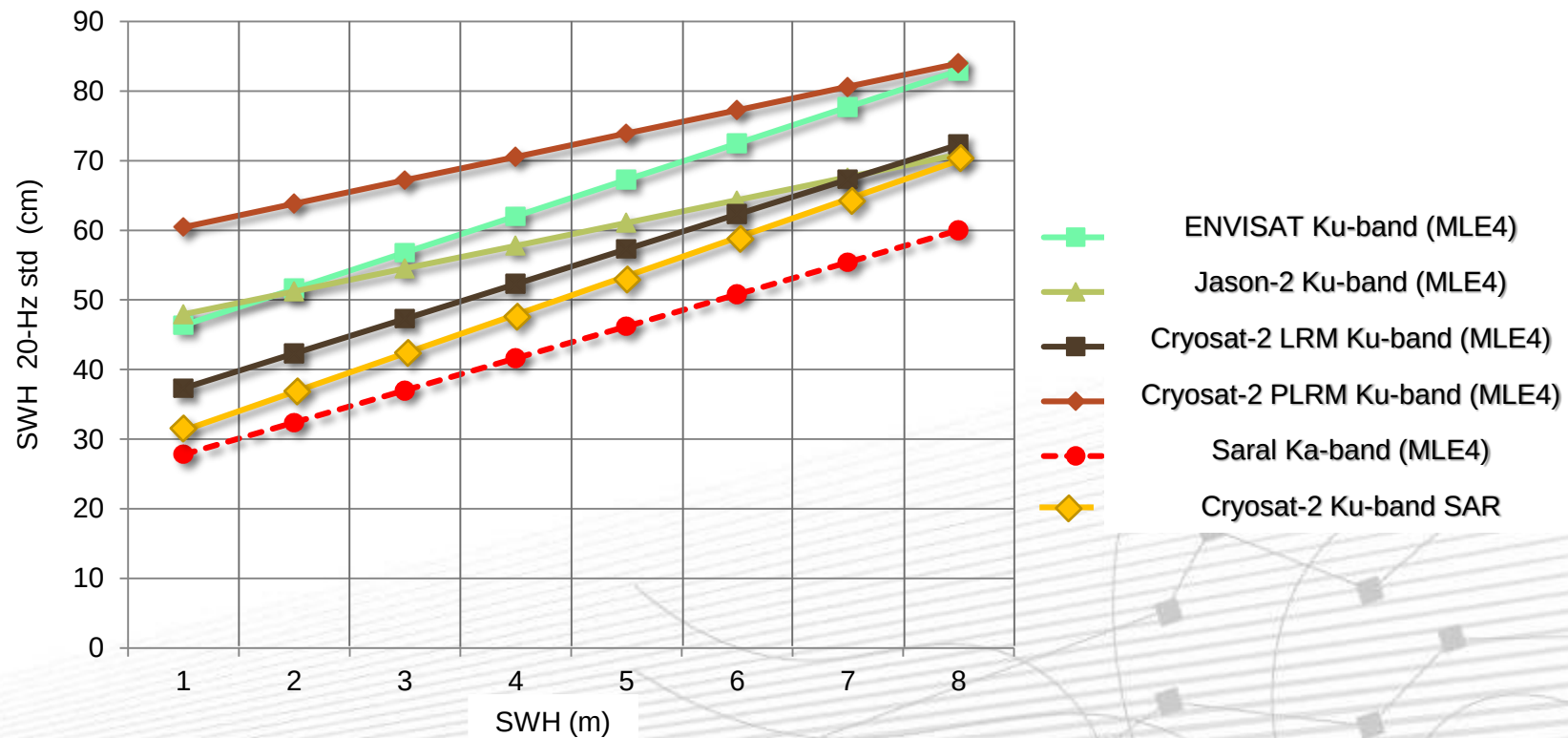
S3A PRLM – S3B LRM



Quite stable above 1m SWH

It can be shown with two different metrics:

- ❑ 20Hz/40Hz standard deviations
- ❑ Power Spectral Density

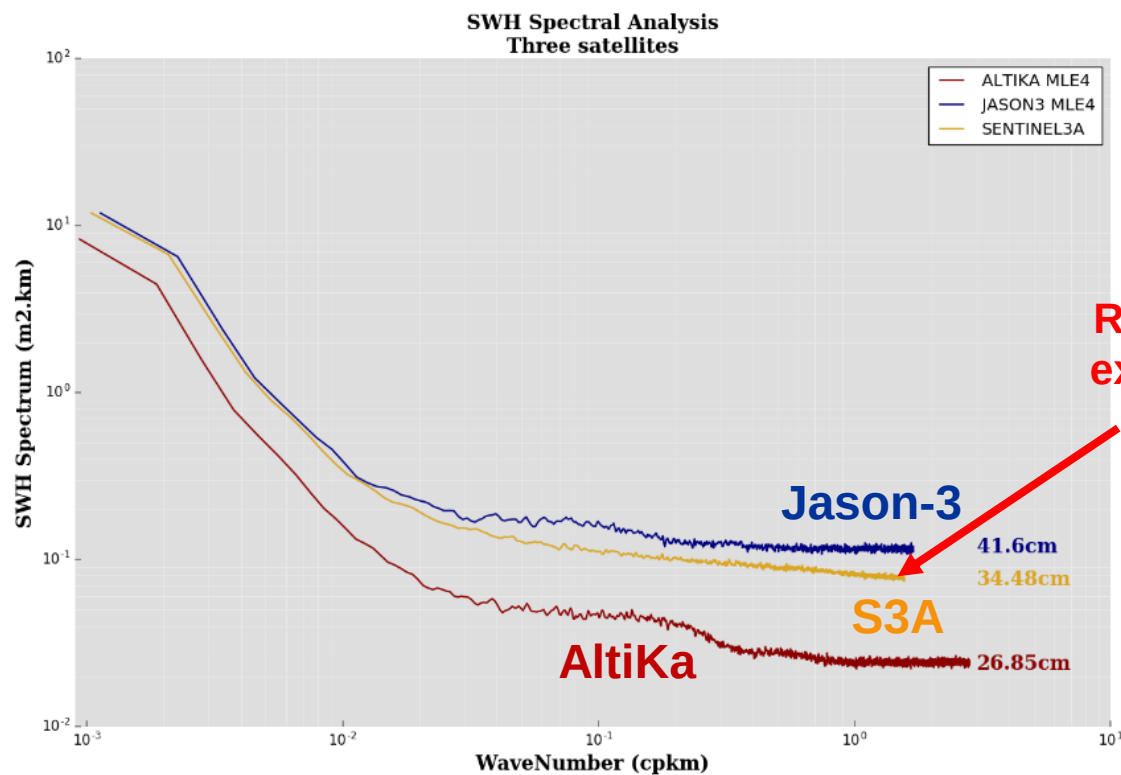


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- ❑ 20Hz/40Hz standard deviations

- ❑ **Power Spectral Density**

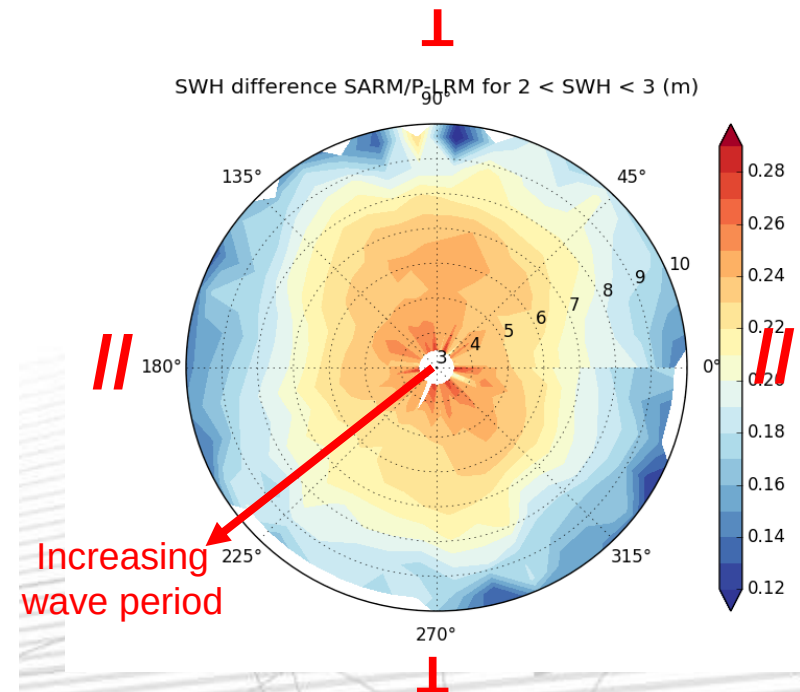
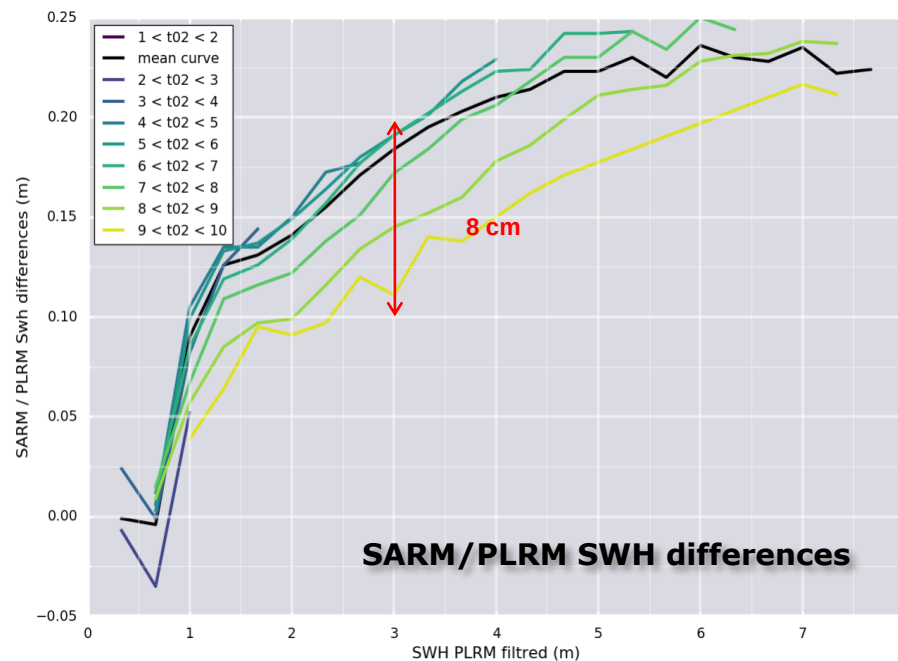
Both metrics are consistent if looking at HF noise level



Red noise in SAR PSD spectrum
explained by the correlation with
swell
(M.Raynal & al, 2018)

[Moreau et al., Adv & Space Research, 2018]

- Increase of noise level with SWH, swell period and for swell parallel to satellite flight direction Strong impact of T02 on SAR mode SWH (8 cm between 5s and 10s period for SWH = 2.7m)



(see P.Thibaut talks in La Rochelle & Miami at OSTST, 2016-2017)

LRM

Cramer Rao Bound computation clearly indicates huge potential to improve SWH estimation (more than other parameters) (Mailhes & al, Proc of Eusipco Conference, Edinburgh, 2008).

The variance of any unbiased estimator of P_u , t and SWH is bounded below by its corresponding Cramer-Rao bound (CRB) which is obtained by inverting the Fischer Information Matrix

CLS/CNES decided to change the estimation method (Newton Raphson in MLE4 but that can be considered as a least square method) and we moved **to a true Maximum Likelihood Estimation method that accounts for speckle distribution** (True Likelihood criterion with a Downhill Simplex method) → Nelder Mead solution

Adaptive Retracker

It accounts as well for the true Instrumental characteristics (PTR) and has been fully validated. 20Hz estimation are directly usable without any Look Up Tables. And many other qualities (for range, σ_0 , peaky echos, continuity,)

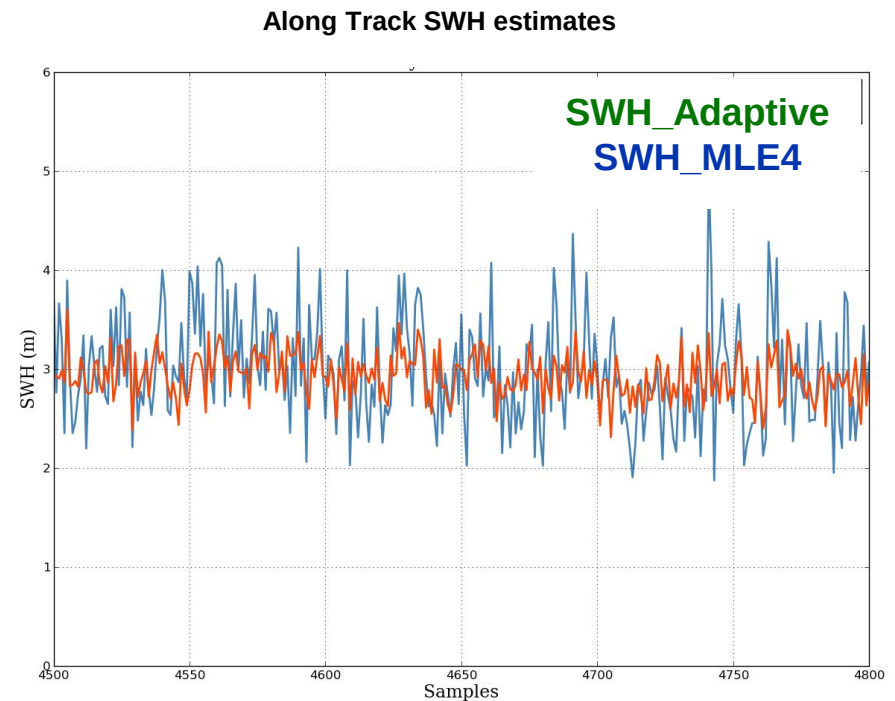
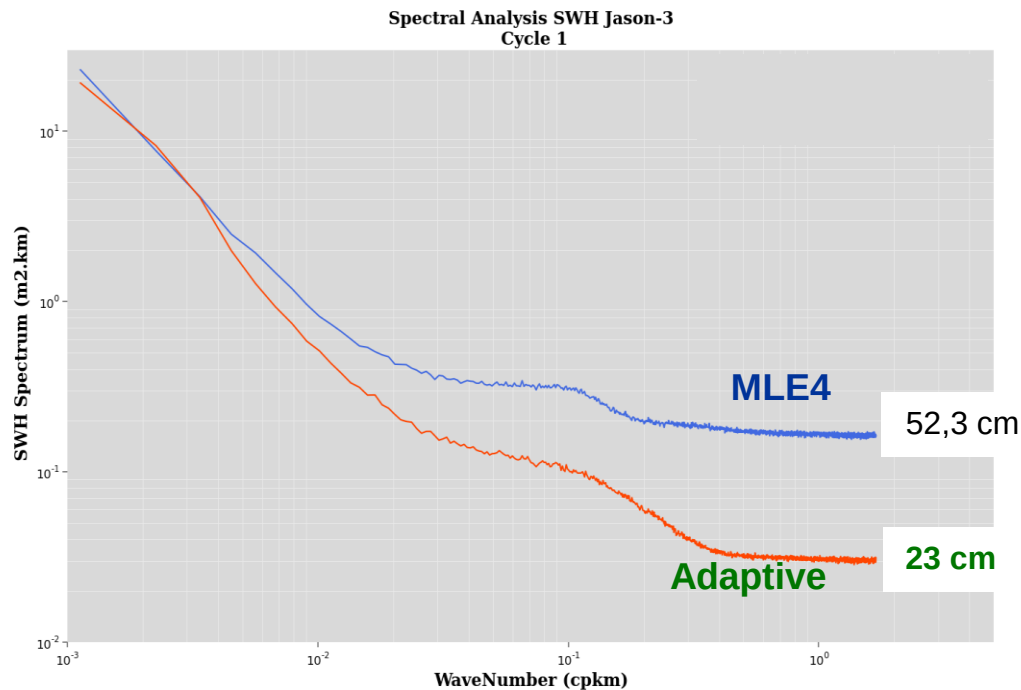
Potential drifts of instrumental features are directly taken into account

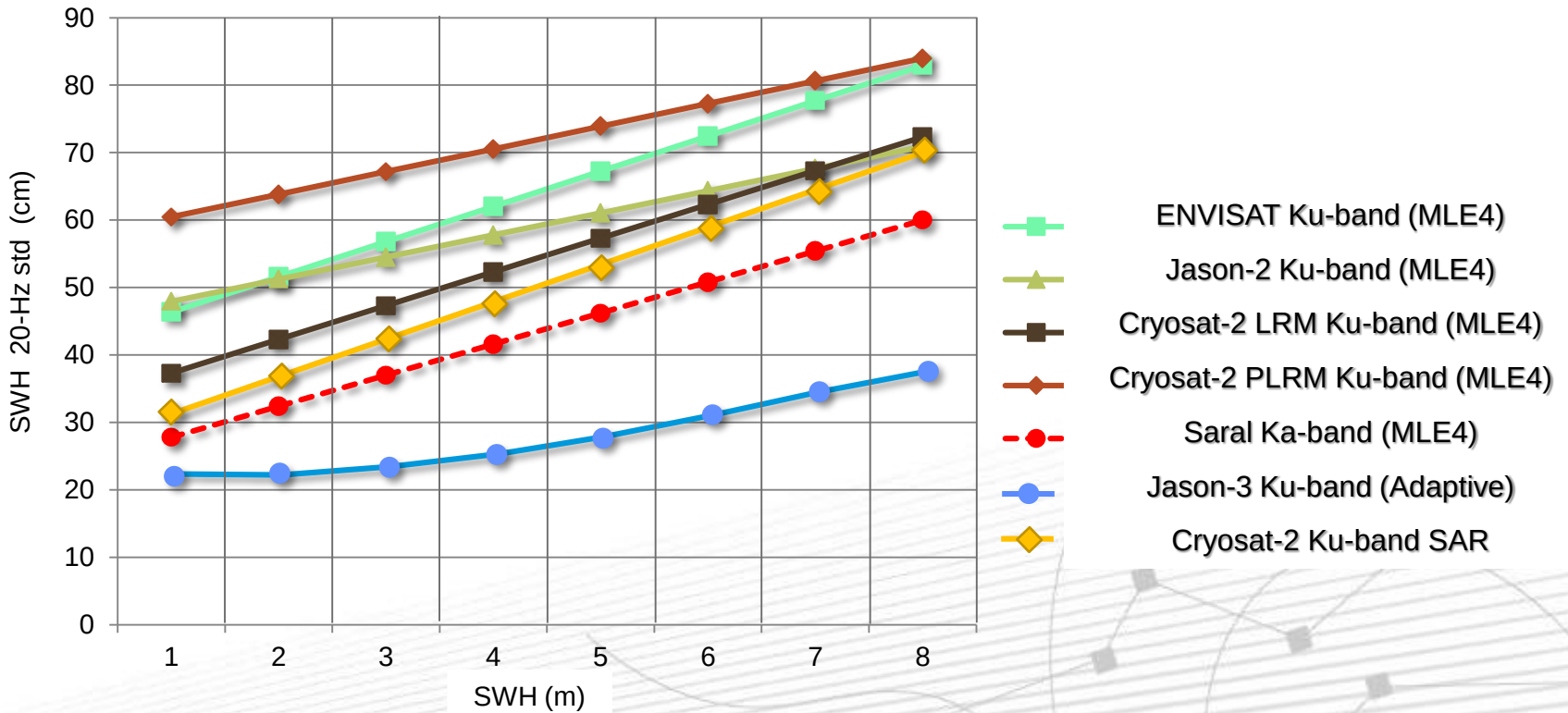
(Note that SAR S3 SWH is drifting by 1 cm/year wrt ECMWF due to evolution of the PTR : S.Dinardo communication to MPC. Not corrected yet)

LRM

Results are spectacular in particular regarding SWH estimates

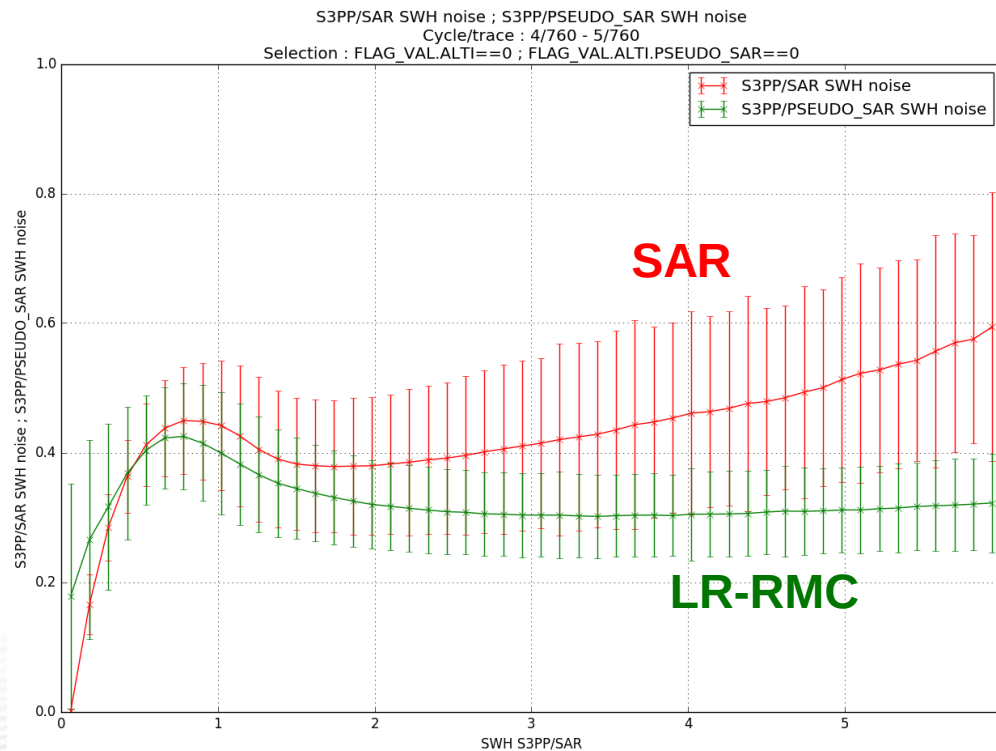
Spectral analysis shows clear reductions of the 20Hz noise level with the Adaptive retracker → -60% for SWH





(see F.Boy talk in Miami OSTST, 2017
& T.Moreau talk, this session)

New way to process the SAR data at Level-1.



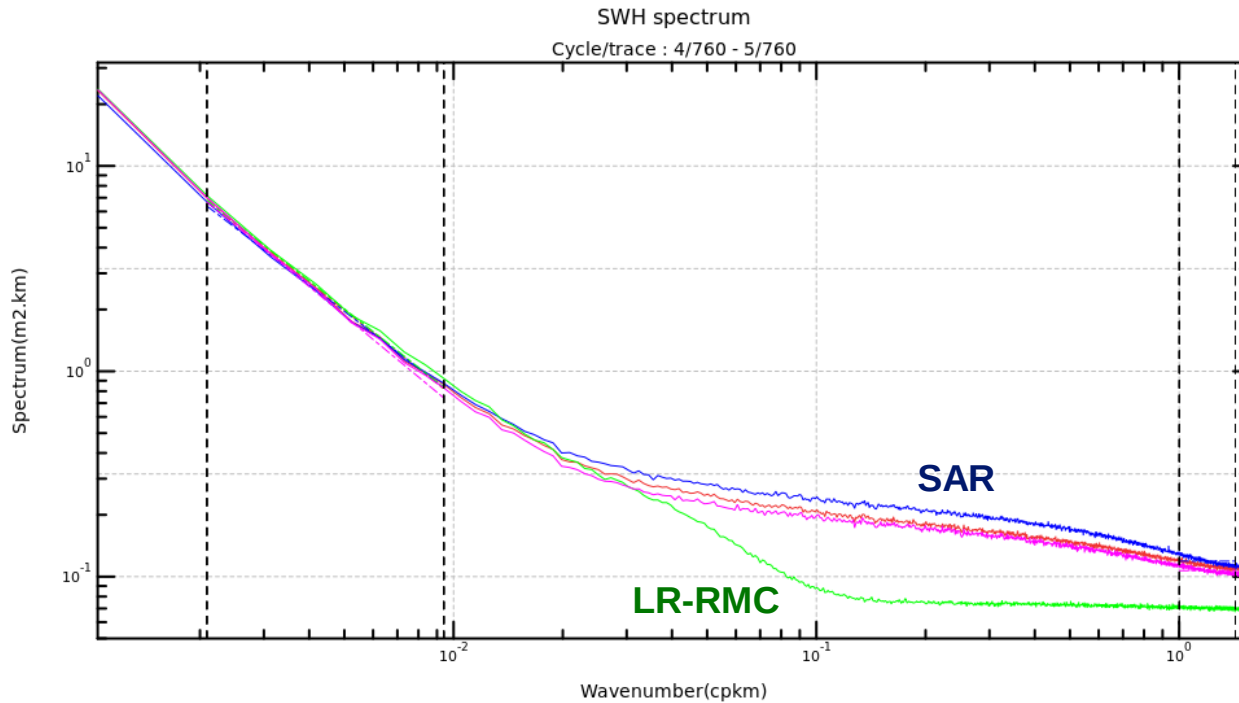
**Clear reduction of
SWH noise level wrt to
SAR (~-40%)**

Solutions

LR-RMC processing (S-3)

(see F.Boy talk in Miami OSTST, 2017
& T.Moreau talk, this session)

SAR



SAR spectrum is largely
impacted by swell
for scales < 50km

LR-RMC spectrum is cleaner
despite a very low filtering
effect at 20km (antenna
footprint)



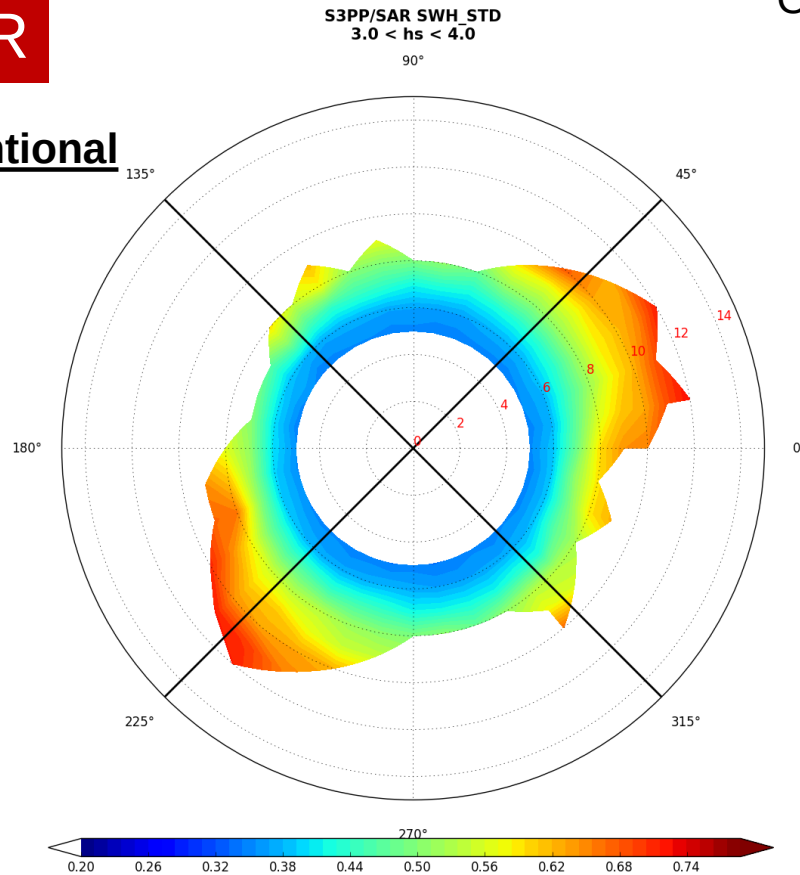
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HAMMING a=-1.35402741447 b=-2.82409799357 sigma=0.431669531412
PSEUDO_SAR a=-1.39714388734 b=-2.91049797361 sigma=0.331972655951
ZERO_PADDING a=-1.46190255917 b=-3.09303721164 sigma=0.409404017942



SAR

On swell induced effects

Conventional
SAR



Noise level depends on sea-state parameters (wavelength and direction)

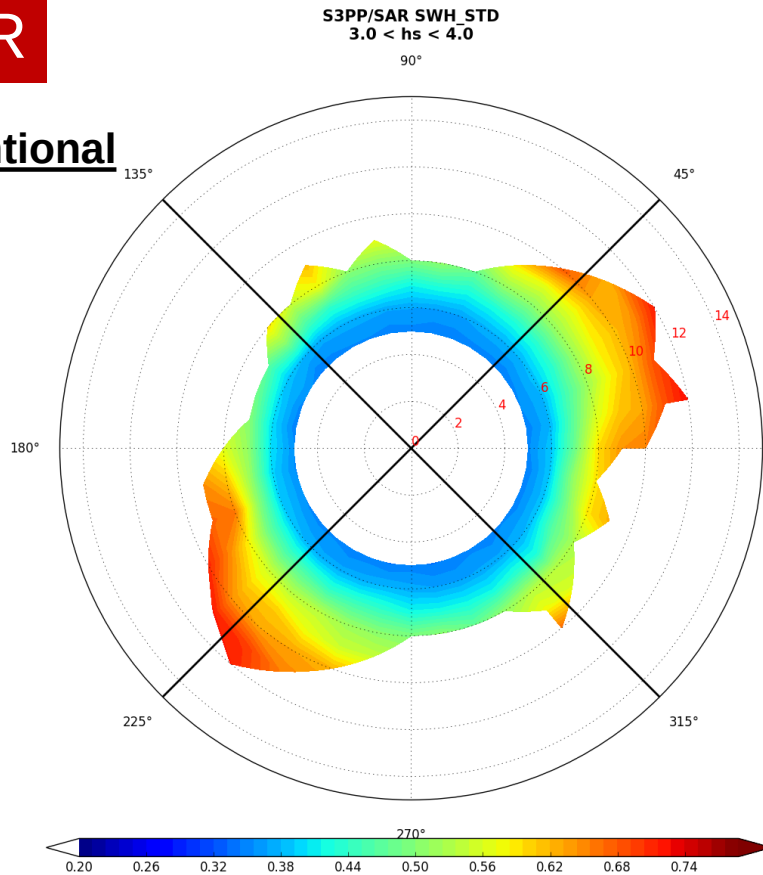
Solutions

LR-RMC processing (S-3)

SAR

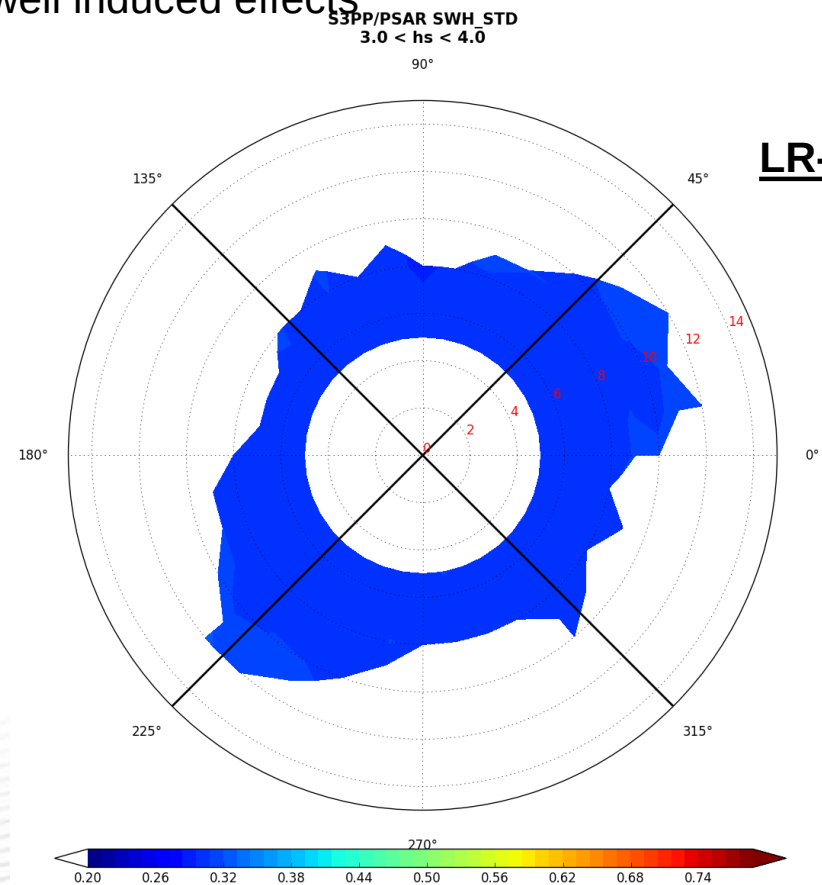
On swell induced effects

Conventional
SAR



Noise level depends on sea-state parameters (wavelength and direction)

LR-RMC



No more dependency on sea-state parameters (wavelength and direction)

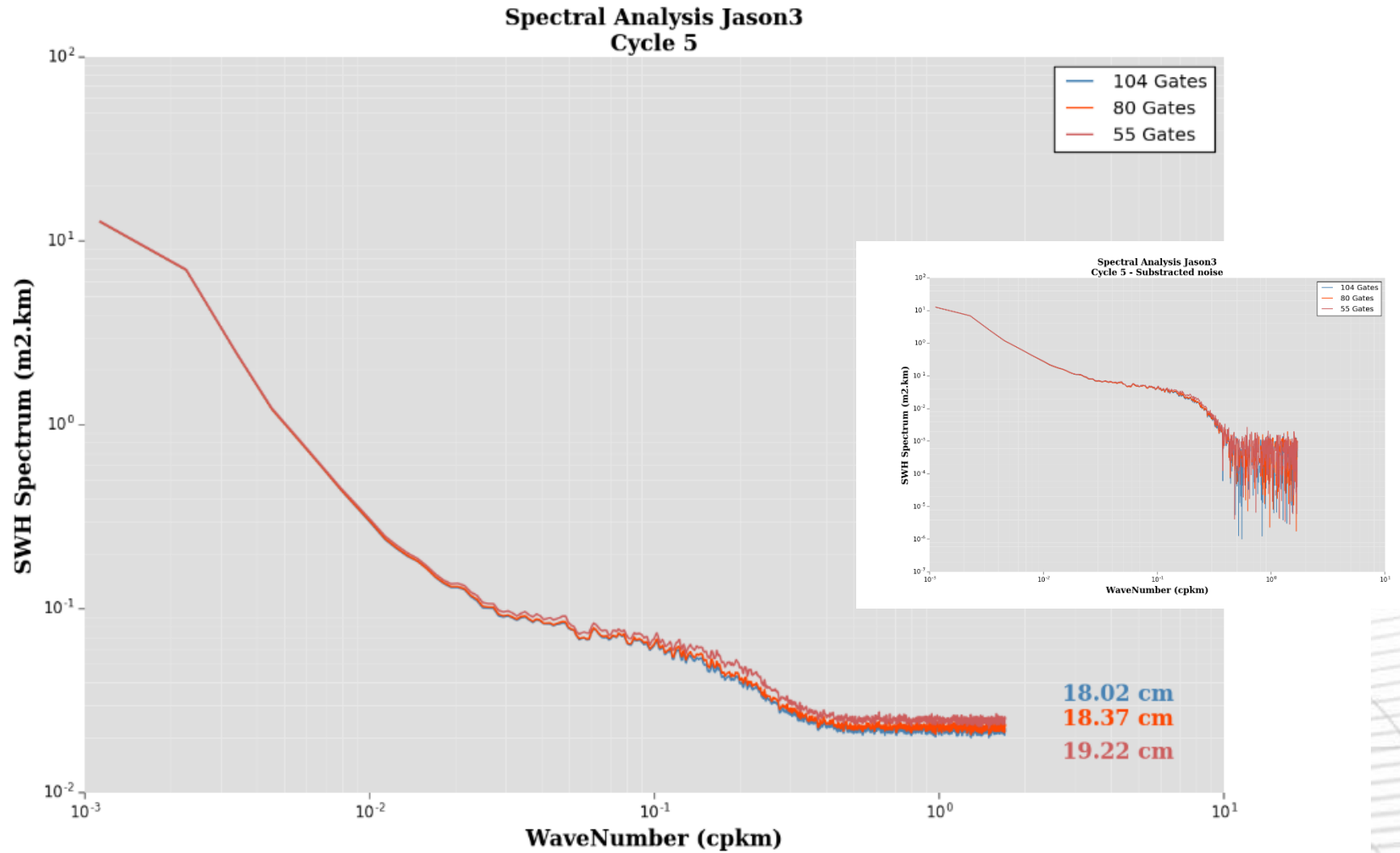
Conclusion

- ❑ Still some improvements to be done on Jason and Saral missions (future reprocessing)
- ❑ Very powerfull «Adaptive method» retracker for LRM ready to be implemented.
- ❑ Swell induced errors observed in SAR can be removed by using LR-RMC L1 processing → clear improvement of performances wrt SAR
- ❑ New retracking strategies oriented towards SWH improvements could be investigated (as 2pass retrackers for SSH, ...)
- ❑ For SAR, performances are still not fully satisfactory and work has to be done before the Sentinel-6 mission which will be a reference mission for Copernicus (Jason-Continuity of Service)
- ❑ CFOSAT mission very soon
- ❑ Ongoing ESA CCI-Sea-State activity conducted by F.Ardhuin to compare different estimation solutions (alti and SAR images) and to produce the best data for each mission. Have a look at M.Passaro poster presenting the algo.Dev team in CCI.

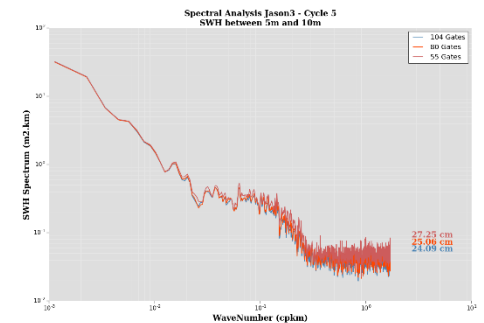
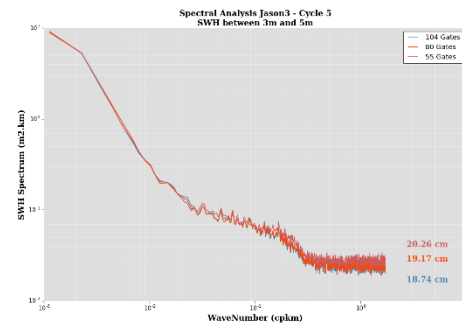
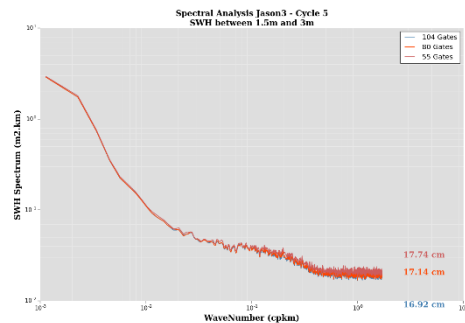
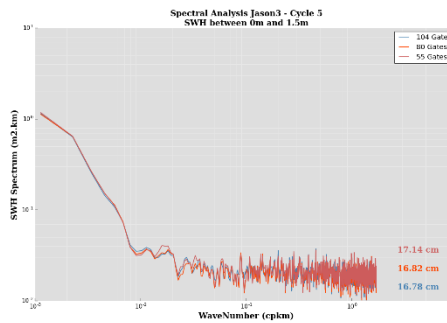
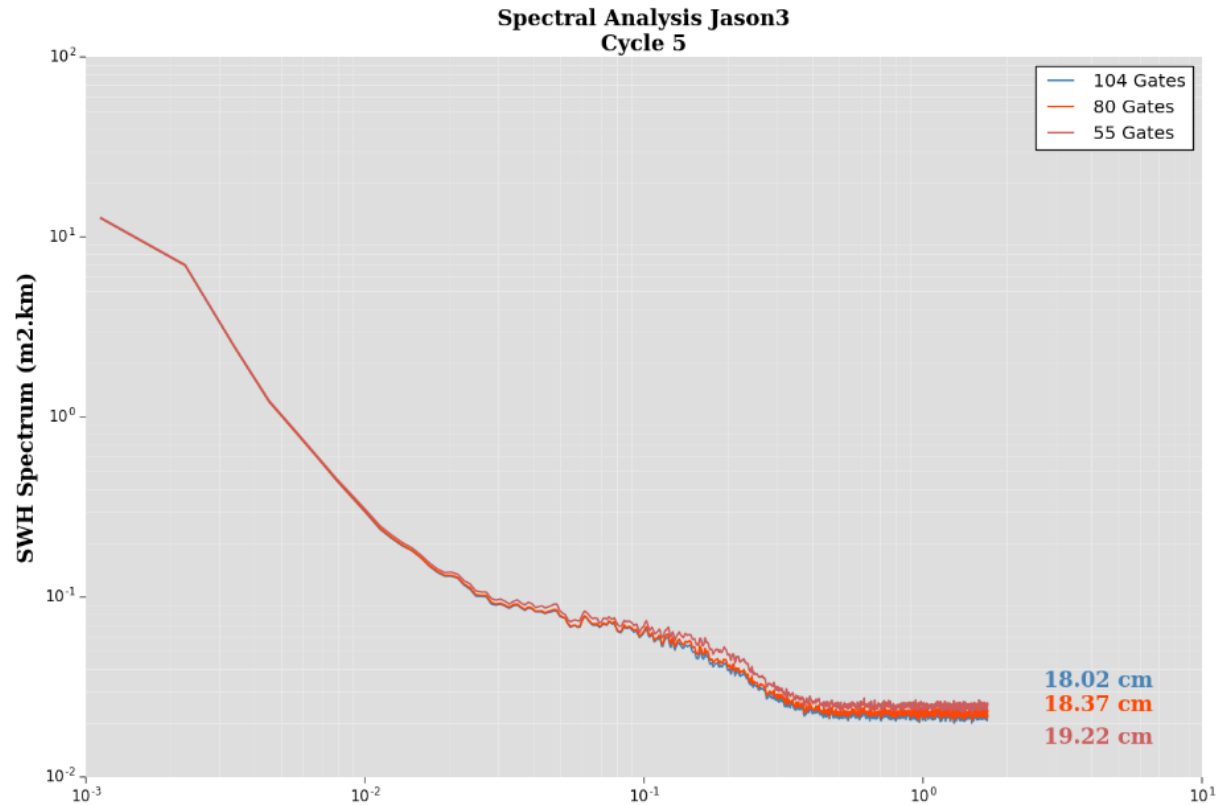
... Thank you for your attention ...

... Back up slides ...

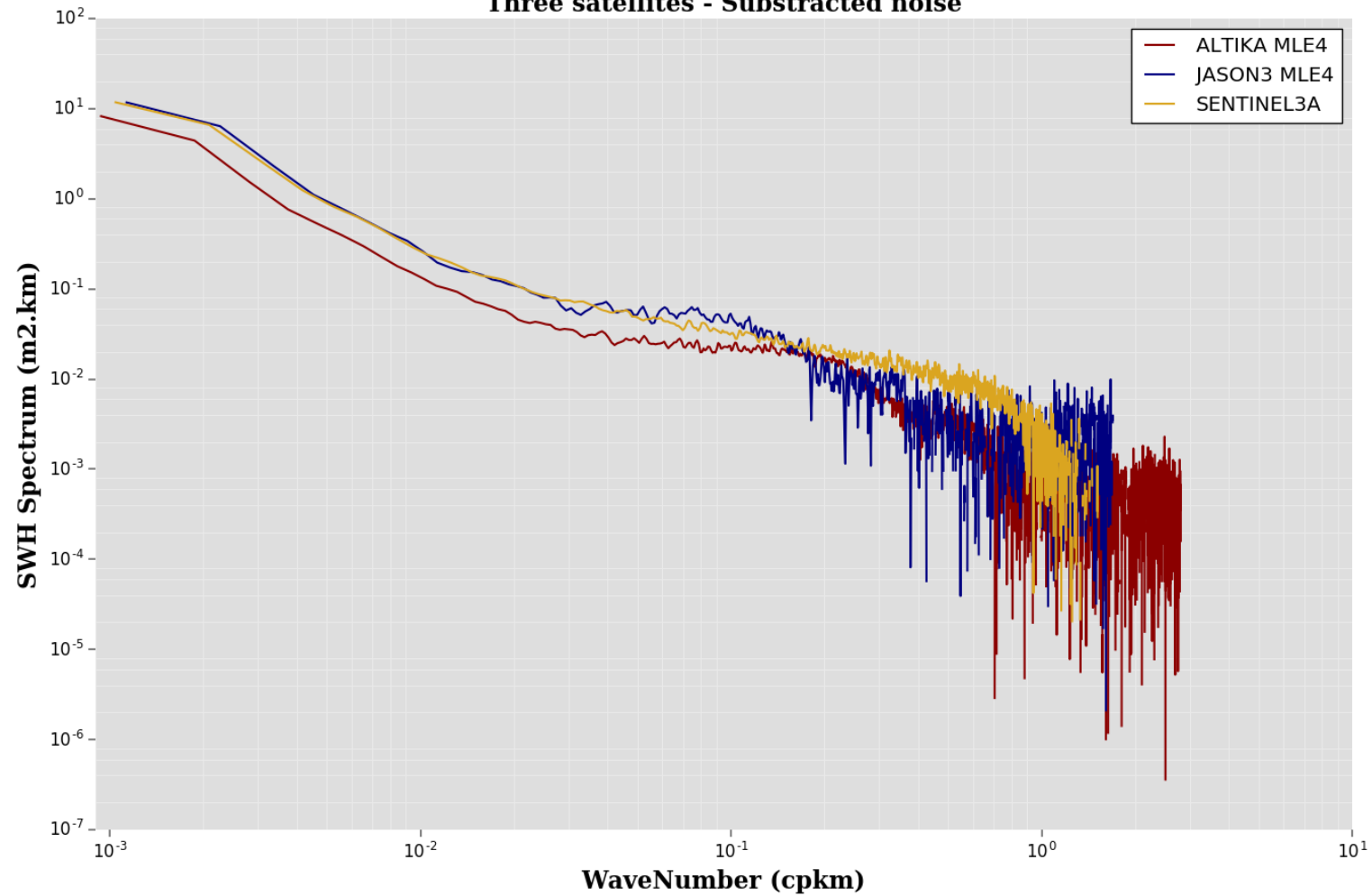
Reduction of the Retracking Window

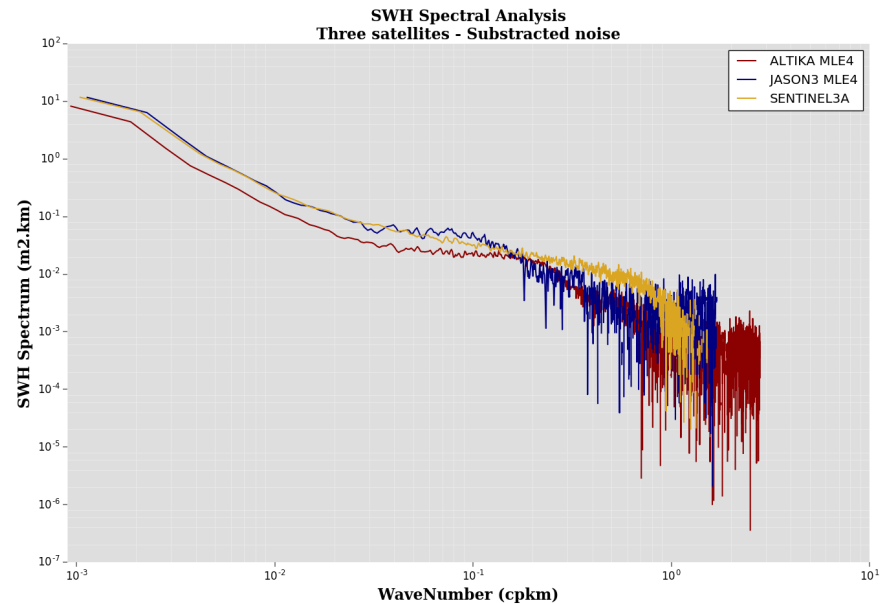
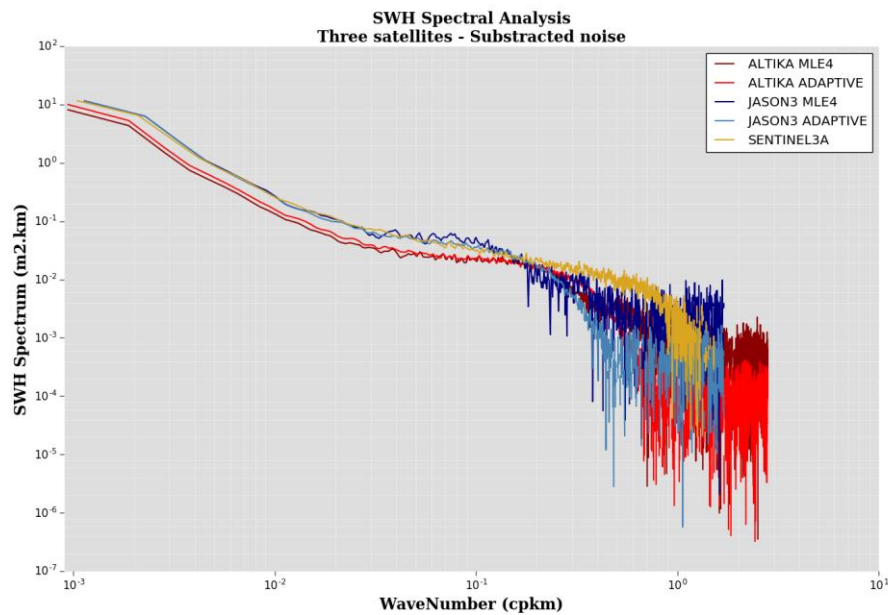


Reduction of the Retracking Window

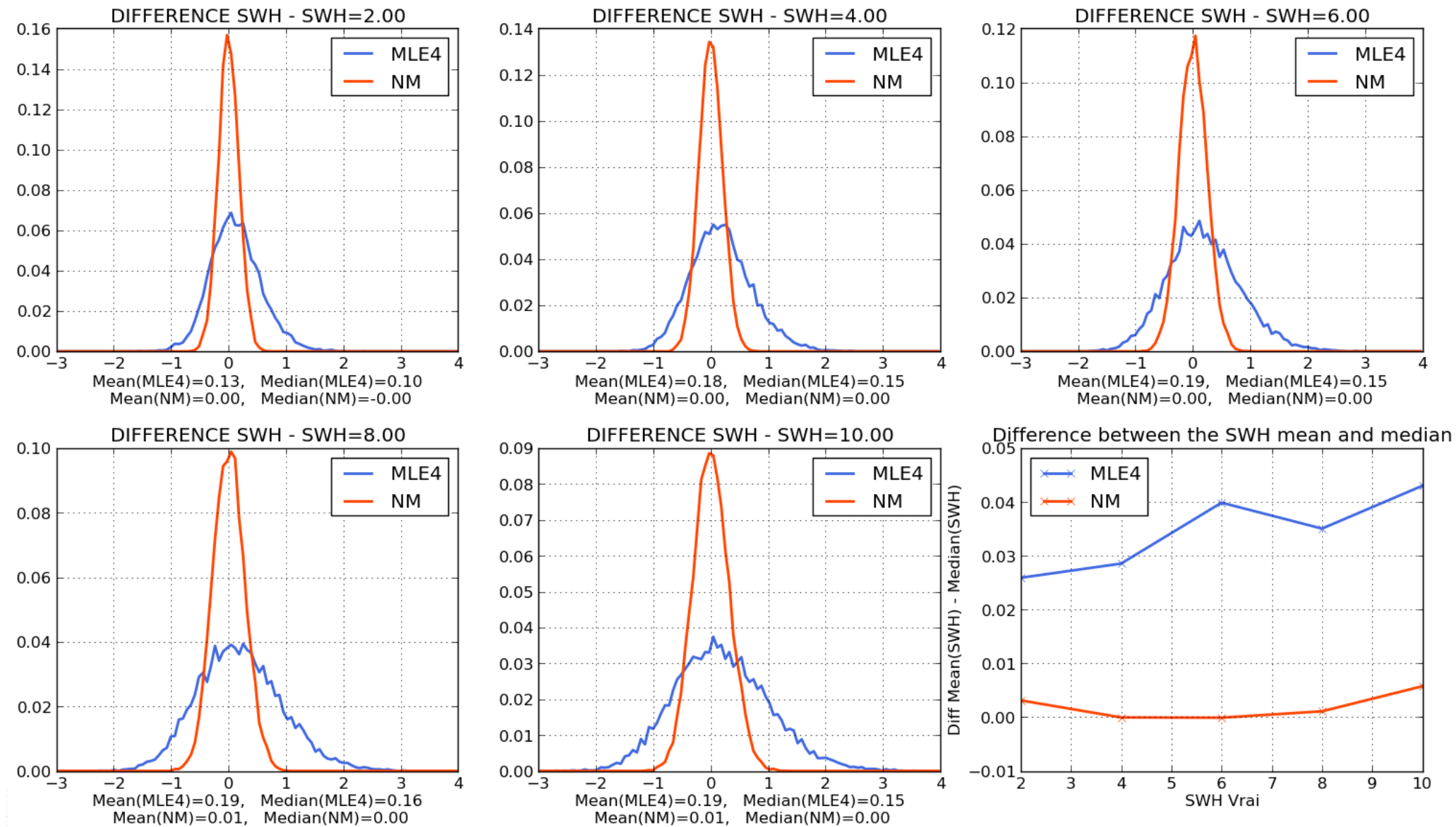


SWH Spectral Analysis Three satellites - Subtracted noise

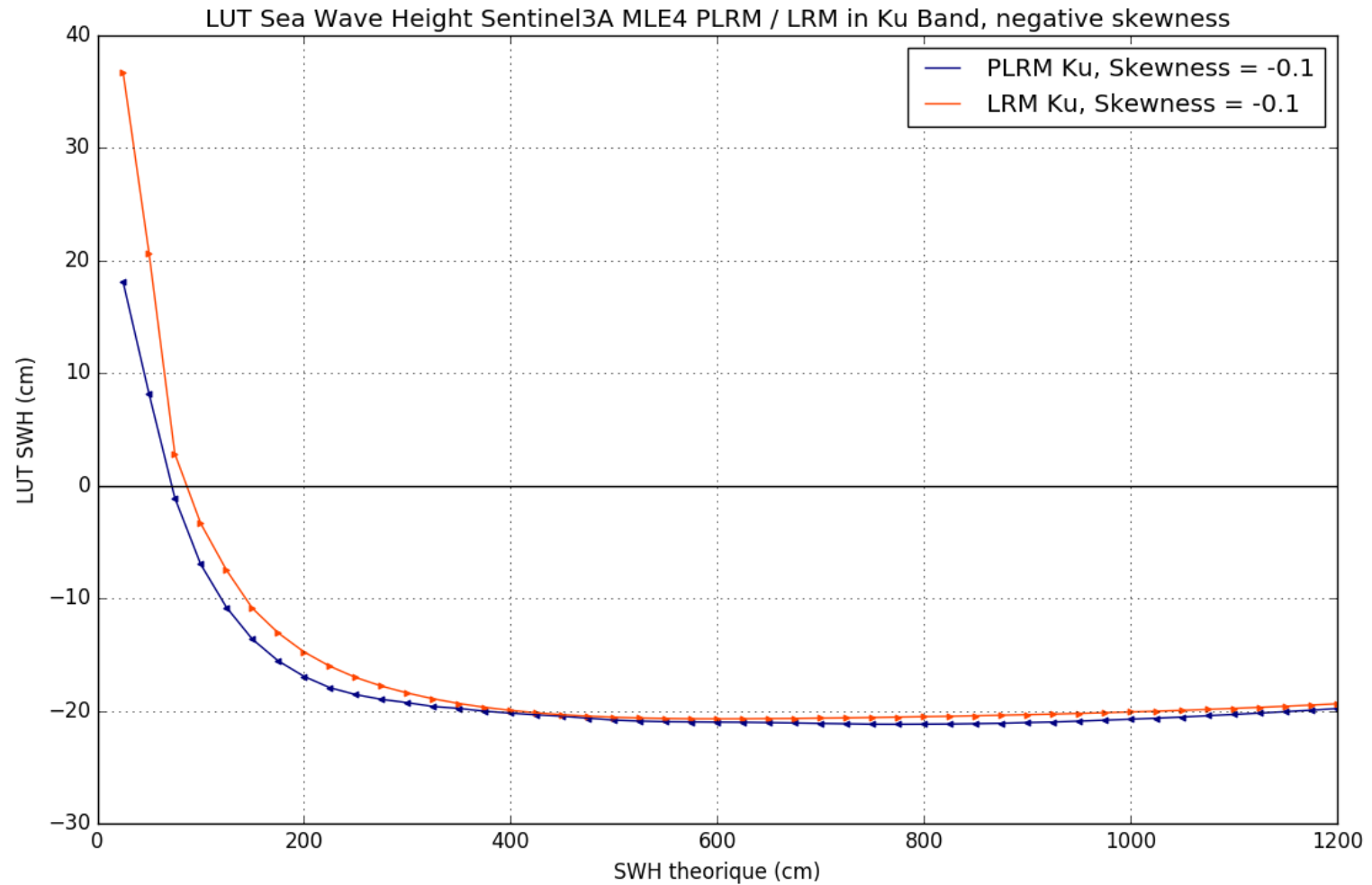




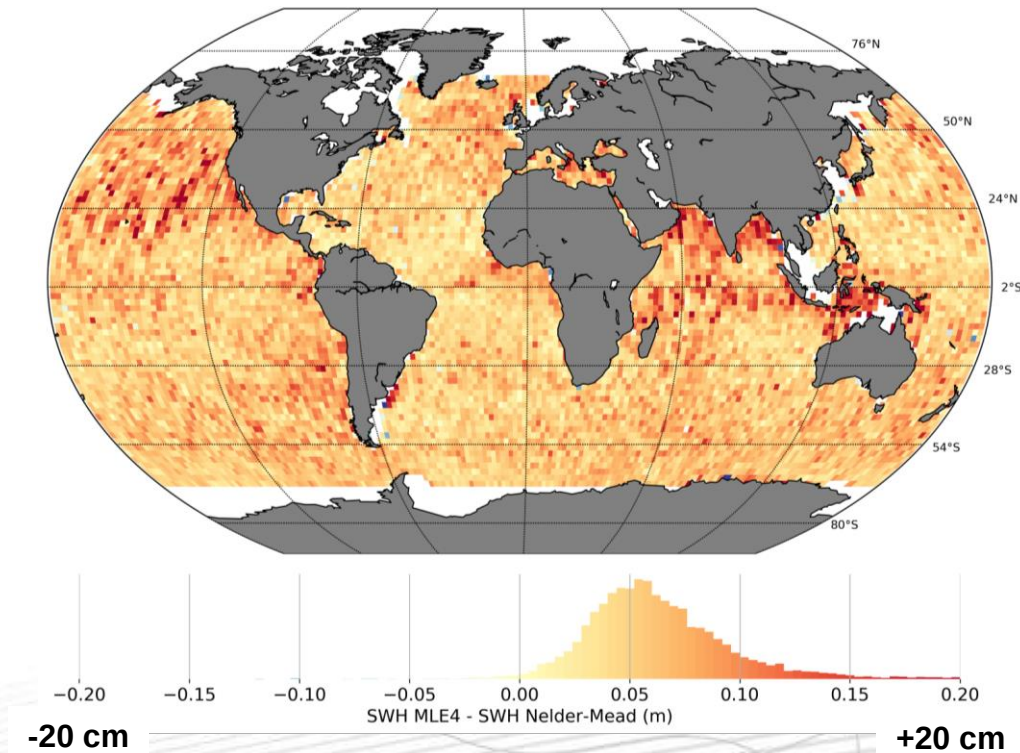
Computation of MLE4 LUTs



Difference between PRL and LRM Look Up Tables for Sentinel-3 A and B



SWH_MLE4 – SWH_Adaptive



*MLE4 estimates
account for LUT*

Comparison of ALTIKA Look-Up Tables
between MLE4 Ku with gaussian approximation and real antenna diagram

